

Supplemental Material for PaperID 0159 “A Reconfigurable Camera Add-On for High Dynamic Range, Multispectral, Polarization, and Light-Field Imaging”

1 This document provides an additional discussion of design choices
 2 and information regarding the processing steps involved in the ap-
 3 plications.

4 1 Details on the Choice of the Diffuser

5 The diffuser structure is of major practical concern for image qual-
 6 ity. Since we are using high resolution sensors, and essentially per-
 7 form a 1:N minification, the diffuser structure should ideally be be-
 8 low $N \times \mu$, where μ is the physical pixel size, to avoid seeing the
 9 diffuser structure in the final image. A large diffusion angle, how-
 10 ever, requires sufficient structures for scattering. This can either be
 11 achieved by larger surface structures (typically 50 – 100 μm) in
 12 relatively thin materials, or by finer scattering structures in thicker
 13 slabs. The former option leads to graininess in the image whereas
 14 the latter induces glare by multiple scattering. We experimented
 15 with both types of diffuser, one thin material that has been extracted
 16 from an LCD screen (thickness $\approx 0.1 mm$), and one diffuser type
 17 that is designed for polarization-based 3D rear projection screens
 18 (thickness $\approx 1 mm$ available from ScreenTech GmbH (material
 19 type “ST-Professional-DCF”). The latter diffuser is designed to be
 20 polarization-preserving. The two diffusers are at opposite ends of
 21 the spectrum, where the LCD diffuser shows considerable graini-
 22 ness, whereas the diffuser intended for 3D rear projection screens
 23 shows considerable glare. Fig. 1 shows the angular scattering pro-
 24 file of both diffusers and measurements of a target that is designed
 25 for the measuring of image resolution and contrast. The LCD dif-
 26 fuser shows graininess but much better contrast than the 3D rear-
 27 projection diffuser which suffers from glare. Unfortunately, the
 28 hotspot nature of the LCD diffuser prevents its use in our filter-
 29 based design since it passes light-field components into the dif-
 30 ferent copies. This means that parallax effects are observable in
 31 the differently-filtered views – clearly an undesirable attribute. We
 32 therefore chose to work with the 3D rear-projection diffuser.

33 2 Details on Multispectral Calibration and Estimation

35 2.1 Spectral Calibration

36 We took images of a MacBeth ColorChecker classic through all
 37 116 broadband filters provided in the Roscolux swatchbook as pro-
 38 duced by Rosco Laboratories. The transmission spectra of these
 39 filters were measured by a spectrometer (CCS 200, Thorlabs
 40 Inc.; 200 nm – 1000 nm). As in the paper we denote them by
 41 f_i , $i = 1 \dots 116$. In addition, we measured the spectral output of
 42 a light source that has a higher efficiency in the blue than an incan-
 43 descent light bulb, a high pressure mercury vapor lamp s_{mv} . The
 44 116 input images are radiometrically compensated. We then ap-
 45 proximate the Bayer filter spectral transmission curves by a linear
 46 combination of basis functions.

$$f^{r|g|b}(\lambda) = \sum_{j=0}^{N-1} a_j^{r|g|b} \phi_j(\lambda). \quad (1)$$

47 Here, N is the number of basis functions ϕ_j used for the approx-
 48 imation and $a_j^{r|g|b}$ are the coefficients for the red, green, and blue
 49 response functions. Inserting into Eq. (3) from the paper, we obtain

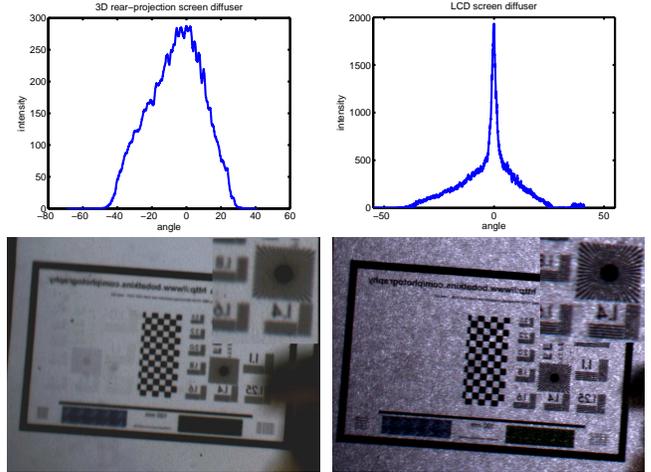


Figure 1: Top Row: Angular response of two different diffusers. Left: A diffuser intended for 3D rear-projection. The scattering profile is broad and does not feature spiky peaks - it is said to be free of hot spots. This type of diffuser is usable in our system. Right: A diffuser extracted from an LCD screen. The diffuser profile shows a strong peak in the center - this peak is called a hot spot in diffuser optical language. It passes light-field components into our imaging system and is unsuitable for filter-based imaging. Bottom Row: Comparison of image quality. Left: Diffuser for 3D rear projection. The contrast is reduced due to glare within the thick material of the diffuser. In this example, the resolution is about 1.2 line pairs per mm. Right: LCD diffuser, the material shows surface structure which has a highly non-uniform angular response. The resolution and contrast, however, are much better than in the 3D rear-projection diffuser, 1.8 line pairs per mm can be resolved. Further, the LCD diffuser results in observable parallax between sub-images.

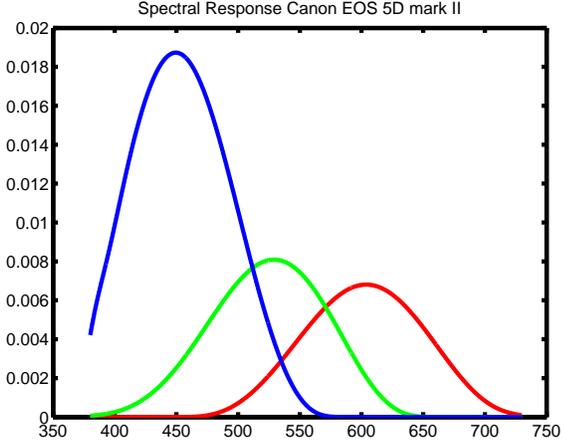


Figure 2: Spectral response of the Canon EOS 5D mark II as estimated from 116 broadband filtered images.

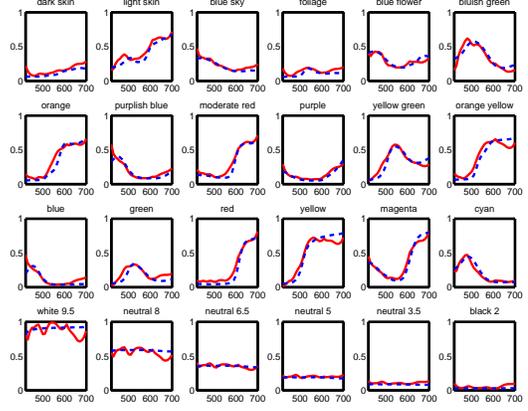


Figure 3: Spectral reflectance of a MacBeth ColorChecker classic as imaged and reconstructed by our system (red) and average ground truth measurements obtained from 30 ColorCheckers (blue), source: http://www.babelcolor.com/main_level/ColorChecker.htm.

a linear system of equations:

$$I_i^{r|g|b}(x, y) = \sum_{j=0}^{N-1} a_j^{r|g|b} \int_{\lambda} l_{\lambda}(x, y, \lambda) \phi_j(\lambda) f_i(\lambda) s_{mv}(\lambda) d\lambda, \quad (2)$$

with the integral now being completely determined. In theory, we could solve the resulting linear systems

$$\mathbf{I}^{r|g|b} = \Phi \mathbf{a}^{r|g|b} \quad (3)$$

in a least-squares sense by inverting $\mathbf{a}^{r|g|b} = (\Phi^T \Phi)^{-1} \mathbf{I}^{r|g|b}$ where the vectors collect the individual coefficients and image values and the matrix $\Phi_{i,j}$ contains the integral values for the i 'th image and the j 'th basis function. In practice, we need to add a standard Laplacian regularization term (with Neuman boundary conditions) and enforce positiveness of the solution. We solve

$$\min_{\mathbf{a}^{r|g|b}} (\Phi \mathbf{a}^{r|g|b} - \mathbf{I}^{r|g|b})^2 + \alpha (\mathbf{L} \mathbf{a}^{r|g|b})^2 \text{ subject to } \mathbf{a}^{r|g|b} \geq 0 \quad (4)$$

using the MATLAB `quadprog` sub-routine. In practice, we use 50 Gaussian basis functions to cover the range between 400 nm and 700 nm, which is the region of best sensitivity of our camera. The regularization parameter α was set to 50. The resulting curves are shown in Fig. 2

2.2 Spectral Image Estimation

Our imaging device creates 9 copies that are filtered with spectral broadband filters selected from the Rosco Labs Roscolux swatchbook. The filters were selected manually, based on spectral coverage considerations. The filters that were used for all experiments are { Cyan #4360, Yellow #4590, Red #26, Orange #23, Green #89, Blue-Green #93, Lavender #4960, Blue #80, Magenta #4760 }. The resulting images are RGB images, i.e. each of the nine filters is modulated by an additional Bayer filter, resulting in 27 measurement channels. We use the same algorithm as for the calibration of the spectral sensitivities, Sec. 2.1. There is one minor modification; following [Toyooka and Hayasaka 1997; Park et al. 2007], we use a PCA basis for the spectral dimension. It is well known, that natural spectral reflectances can be modelled by a low-dimensional linear model [Jaaskelainen et al. 1990]. This modification necessitates an adaptation of the Laplacian prior since regularization is to be performed on the reconstructed function, not on the coefficients

of its linear expansion. For this purpose, an additional matrix Ψ containing the basis vectors in its columns is introduced into the regularization term.

$$\min_{\mathbf{a}^{r|g|b}} (\Phi \mathbf{a}^{r|g|b} - \mathbf{I}^{r|g|b})^2 + \alpha (\mathbf{L} \Psi \mathbf{a}^{r|g|b})^2. \quad (5)$$

Moreover, since a large number of pixels have to be reconstructed, we cannot afford to solve a quadratic program in this case. We therefore resort to solving the standard per-pixel least-squares problem

$$\mathbf{a}^{r|g|b} = (\Phi^T \Phi + \alpha (\mathbf{L} \Psi)^T (\mathbf{L} \Psi))^{-1} \mathbf{I}^{r|g|b}. \quad (6)$$

We construct the PCA basis by analyzing a database of 1269 spectra measured from the Munsell book of colors (source: <http://www.uef.fi/spectral/munsell-colors-matt-spectrofotometer-measured>) in agreement with [Jaaskelainen et al. 1990]. In practice, we use the mean of the data set and the first 14 PCA vectors as basis functions. The regularization parameter α has been set to a low value of 0.01 since the basis itself has strong regularizing properties.

Evaluation To evaluate our reconstruction scheme, we recorded a MacBeth Color checker classic under the illumination of a high-pressure mercury vapor lamp. The results are compared to average measurements obtained from spectrometers that were collected on http://www.babelcolor.com/main_level/ColorChecker.htm, see Fig. 3. The results are in good agreement with the collected data.

2.3 Relighting and Simulation of Color-Deficiency

Our applications are based on a modulation of the acquired color spectra. The processing chain involves the computation of neutral spectral images with 29 wavebands. Since our examples have been recorded in a lab setting, the spectral composition of the illuminating source was known (measured by spectrometer Thorlabs CCS 200). We can then modulate the spectral image stack with illuminating spectra from different sources:

- a simulation of different black body spectra using Planck's law,

- 113 • an assortment of daylight spectra obtained from <http://www.uef.fi/spectral/daylight-spectra>,
- 114
- 115 • an assortment of artificial light spectra obtained from <http://www.uef.fi/spectral/artificial-lights>, and
- 116
- 117
- 118 • a simulated sweeping band pass with Gaussian profile of
- 119 20 nm width.

120 The resulting spectral images were converted to XYZ using the
 121 CIE64 standard observer. Finally, a conversion to sRGB (gamma
 122 2.2) with a simulated D65 daylight source was performed.

123 **Simulation of Color Deficiency** For simulating the vision of
 124 color deficient subjects, we obtained the color-matching functions
 125 for protanomalous (response of the red cones shifts towards green
 126 by about 10 nm) and deutanomalous subjects (response of the
 127 green cones shifts towards red by about 10 nm) [DeMarco et al.
 128 1992]. These two are the most common color deficiencies en-
 129 countered in the population. In addition, there are three types
 130 (protanopia, deutanopia, tritanopia) that have zero response for
 131 the L, M, and S cones, respectively. The simulations were per-
 132 formed by substituting the CIE64 color matching functions with
 133 their adapted versions. Since these are not designed to convert to
 134 XYZ but to an LMS cone response, we first mapped the output to
 135 XYZ, which was then converted to sRGB for display.

136 3 Detailed Derivation of the Horn-Schunck- 137 based Depth Estimation

138 To derive a depth map, we do not rely on stereo matching, but in-
 139 stead employ a modified optical flow algorithm. We adapted the
 140 Horn and Schunck functional by introducing a scalar depth func-
 141 tion $d(x, y)$ for the center view. This depth value results in a paral-
 142 lax displacement in the neighboring images via $d \cdot [u_i, v_i]^T$, where
 143 the vector $[u_i, v_i]^T$ denotes the direction of the epipolar lines in a
 144 neighboring view I_i . Our goal is to estimate a depth that best ex-
 145 plains all views in a least-squares sense. The resulting error func-
 146 tional adapted from Horn & Schunck flow is given by

$$\min_d E = \int \sum_i (d(\nabla I_i \cdot [u_i, v_i]^T) + I_i^t)^2 + \alpha^2 \|\nabla d\|^2 dx dy, \quad (7)$$

147 where i is an index denoting the neighboring views, excluding the
 148 center. $[u_i, v_i]$ are constant vectors that indicate the direction of
 149 the epipolar lines in view i , $\nabla I_i = [I_i^x, I_i^y]$ is the spatial gradient
 150 of view i , and I_i^t denotes the “temporal” derivative between view
 151 i and the center view. The term $\alpha^2 \|\nabla d\|^2$ is a standard Laplacian
 152 smoothness term on the depth map with regularizing parameter α^2 .
 153 The scalar function $d(x, y)$ is the quantity being optimized for. The
 154 corresponding Euler-Lagrange equation is

$$d \cdot \sum_i (I_i^x u_i + I_i^y v_i)^2 + \sum_i I_i^t (I_i^x u_i + I_i^y v_i) - \alpha^2 \Delta d = 0. \quad (8)$$

155 We solve it by standard techniques, discretizing the spatial deriva-
 156 tives by central differences with Neumann boundary conditions and
 157 the “temporal” derivative by Euler forward differencing. We use the
 158 Horn & Schunck approximation to the Laplacian $\Delta d = (d - \bar{d})$,
 159 where

$$\bar{d} = \frac{1}{12} (d_{x-1,y-1} + d_{x+1,y-1} + d_{x-1,y+1} + d_{x+1,y+1}) + \frac{1}{6} (d_{x,y-1} + d_{x,y+1} + d_{x-1,y} + d_{x+1,y}).$$

The equation

$$d^{k+1} = \frac{\alpha^2 \bar{d}^k - \sum_i I_i^t (I_i^x u_i + I_i^y v_i)}{(I_i^x u_i + I_i^y v_i)^2 + \alpha^2}$$

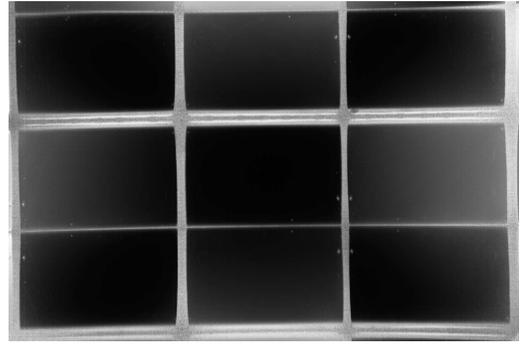


Figure 4: Demonstration of how the optical system changes the polarization state of the incoming light. The image shows the distribution of degree of polarization in the camera’s red color channel. The coding uses black to indicate totally polarized (100%) and white to show unpolarized (0%) states.

160 then defines an update rule to solve for d by Jacobi iterations
 161 k . The scheme is implemented in a standard scale-space fash-
 162 ion [Meinhardt-Llopis and Prez 2012] to allow for large displace-
 163 ments.

164 4 Details on the Requirements and Calibra- 165 tion for Polarization Imaging

166 To allow for polarization imaging with our optical system, it is es-
 167 sential that we employ a polarization-preserving diffuser instead of
 168 a regular diffuser, since a regular diffuser essentially acts as a de-
 169 polarizer. Light traversing through a depolarizer becomes unpolar-
 170 ized regardless of its initial polarization state. In other terms, the
 171 Mueller-matrix of a depolarizer is not a full row rank matrix and,
 172 therefore, retrieving the original Stokes vector of the light becomes
 173 impossible.

174 The other prerequisite is that we need to determine an effective
 175 Mueller-matrix $M_{\text{sys}}(x, y)$ of the system for each pixel (x, y) . It
 176 characterizes the optical system in the sense of how it changes the
 177 polarization state of the incoming light. In Fig. 4, we demonstrate
 178 this effect in an example image, which shows the measured linear
 179 degree of polarization p_{deg} across the nine sub-images in an optical
 180 system without any polarization filters. By definition, we can de-
 181 rive the degree of polarization $p_{deg} = \sqrt{(s^{(1)})^2 + (s^{(2)})^2} / s^{(0)}$ from
 182 the Stokes vector $\mathbf{s} = [s^{(0)} \ s^{(1)} \ s^{(2)}]$. In this particular case, we let
 183 a diffused and completely polarized light beam through the setup.
 184 While at the center, the system retained the 100% degree of polar-
 185 ization, towards the edges the light became less polarized due to the
 186 reflections on the mirrors. We need to compensate for this change
 187 by carrying out a calibration procedure with the complete system
 188 using differently oriented polarizers in the filter array.

189 Basically, the calibration is a Mueller matrix polarimetry whose
 190 principles and possible realization scenarios are thoroughly de-
 191 scribed in [Goldstein 2003]. It needs a system with a complete
 192 polarization state generator (PSG) and a complete polarization state
 193 analyzer (PSA) component. For PSG, we have used an incandes-
 194 cent light source, a paper diffuser, and a polarizer that were placed
 195 at the entrance of the optical system. With these simple tools both
 196 states (totally linearly polarized and unpolarized) can be generated.
 197 We place these elements in the following order; for totally polarized
 198 light *illumination* \rightarrow *paper diffuser* \rightarrow *polarizer* \rightarrow *optical system*
 199 and for unpolarized light *illumination* \rightarrow *polarizer* \rightarrow *optical sys-*
 200 *tem*. Here, we exploit the feature of the paper diffuser, which acts

201 as an almost ideal depolarizer. Further, by still retaining the polar-
 202 izer in the optical path, we can ensure that the light beam entering
 203 the optical system is going to have the same spectrum regardless
 204 of the generated polarization state. With various orientations of the
 205 polarizer, five different totally polarized states and one unpolarized
 206 state were reproducibly created. These states were measured by a
 207 ground-truth imaging polarimeter consisting of a camera, lens (the
 208 same as in our optical system) and a manually rotated polarizer
 209 (similar to [Neumann et al. 2008]). Ground-truth Stokes vectors
 210 $\mathbf{s}_{gt}^{(i)}$, $i = 1..6$, were determined by averaging the obtained Stokes-
 211 vectors within a window of 1500×1500 pixels in the center of
 212 the images, which were completely filled by the unfocused image
 213 of the diffused light beam. Then the same measurements of these
 214 states were carried out through the optical system, employing the
 215 same camera, lens, and the additional manually-rotated polarizer,
 216 yielding a Stokes vector $\mathbf{s}_{sys}^{(i)}(x, y)$ for each pixel (x, y) . These
 217 Stokes-vectors are then matched against to the ground truth ones
 218 using the linear relation $\mathbf{s}_{sys}^{(i)}(x, y) = \mathbf{M}_{sys}(x, y)\mathbf{s}_{gt}^{(i)}$, $i = 1..6$.
 219 $\mathbf{M}_{sys}(x, y)$ is solved via a least-squares regression. Finally, the
 220 obtained system Mueller matrices $\mathbf{M}_{sys}(x, y)$ have to be registered
 221 with the same transformation that is applied to the sub-images and
 222 will yield five Mueller-matrices \mathbf{M}_i , $i = 1..5$ in each sub-image
 223 pixel. These matrices are used for retrieving pixel-by-pixel Stokes-
 224 vectors of the optical system itself and we can follow the description
 225 given in the main paper (see Eq.(4)).

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